

# Vacuum Support for a Large Interferometric Reference Surface

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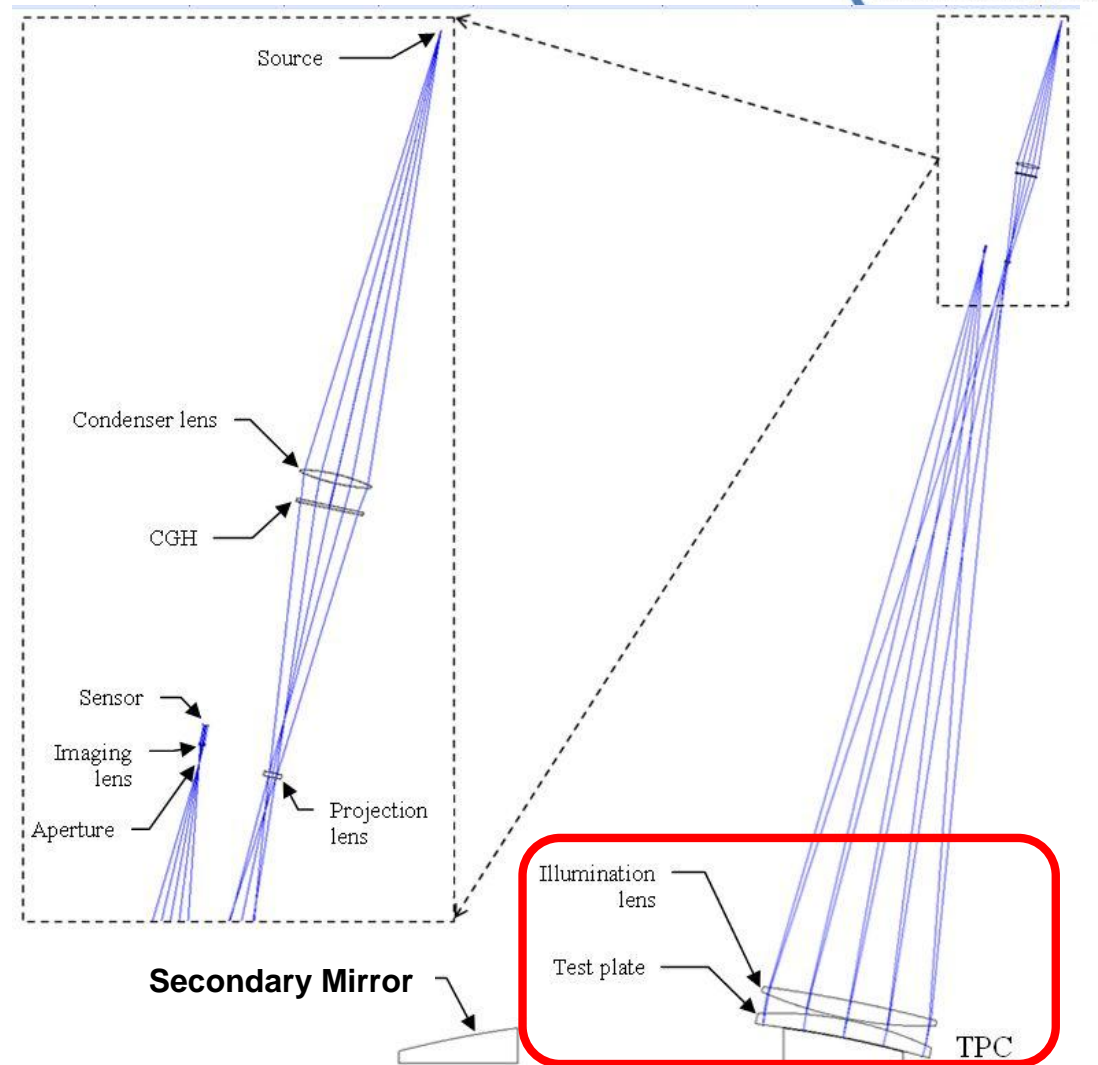
# Outline



- Background
- Purpose
- Requirements for the system
- Experiments for a simple model
- Simulations for an actual model
- Summary

# Background

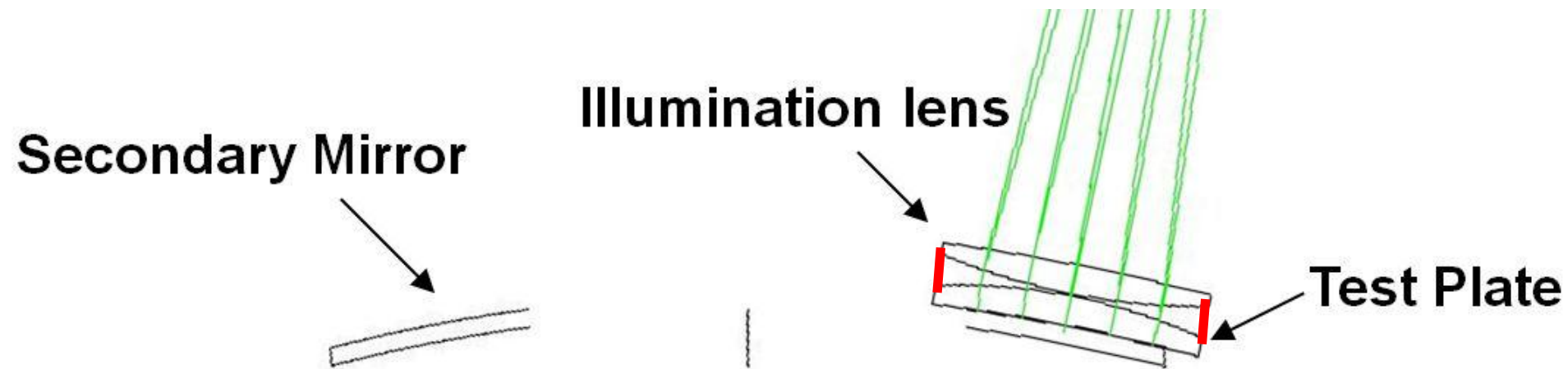
- We try to measure a secondary mirror for a telescope by Fizeau type interferometer with 5mm gap as shown in the right Figure [1].
- In this case, the bending of the test plate caused by the gravity becomes a problem, since the test plate is flipped around after polishing the reference surface.
- Reducing an unexpected bending in the reference surface is required.



[1] M. B. Dubin, et al., "Fizeau interferometer with spherical reference and CGH correction for measuring large convex aspheres", Proceedings of the SPIE, Volume 7426 (2009)

# Purpose

- The scope of this study is to show an effectiveness of the vacuum support for the large interferometric reference surface by simulations.
  - Validate simulations by simple experiments.
  - A mechanical mount for the test plate is also discussed.

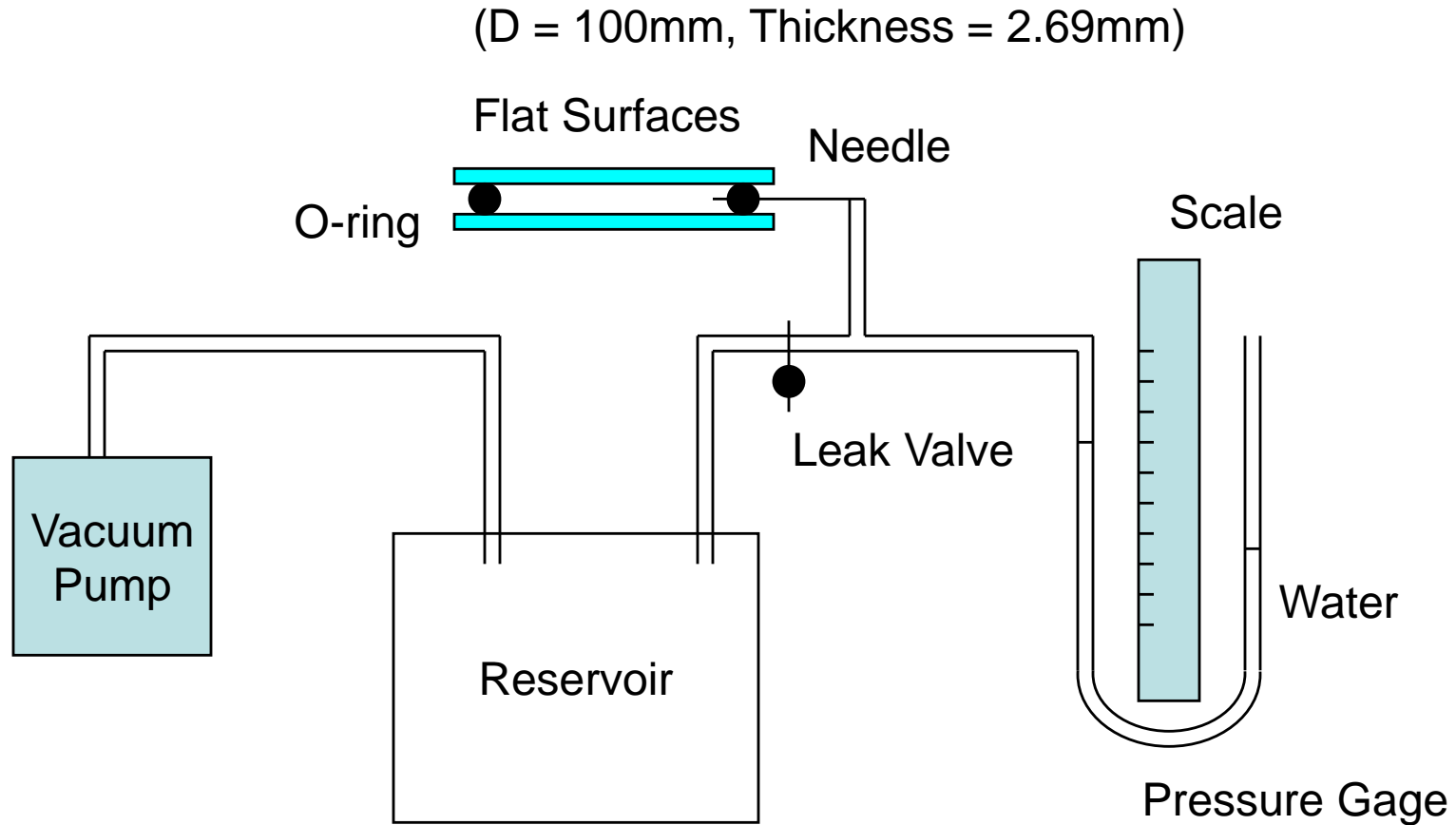


# Requirements



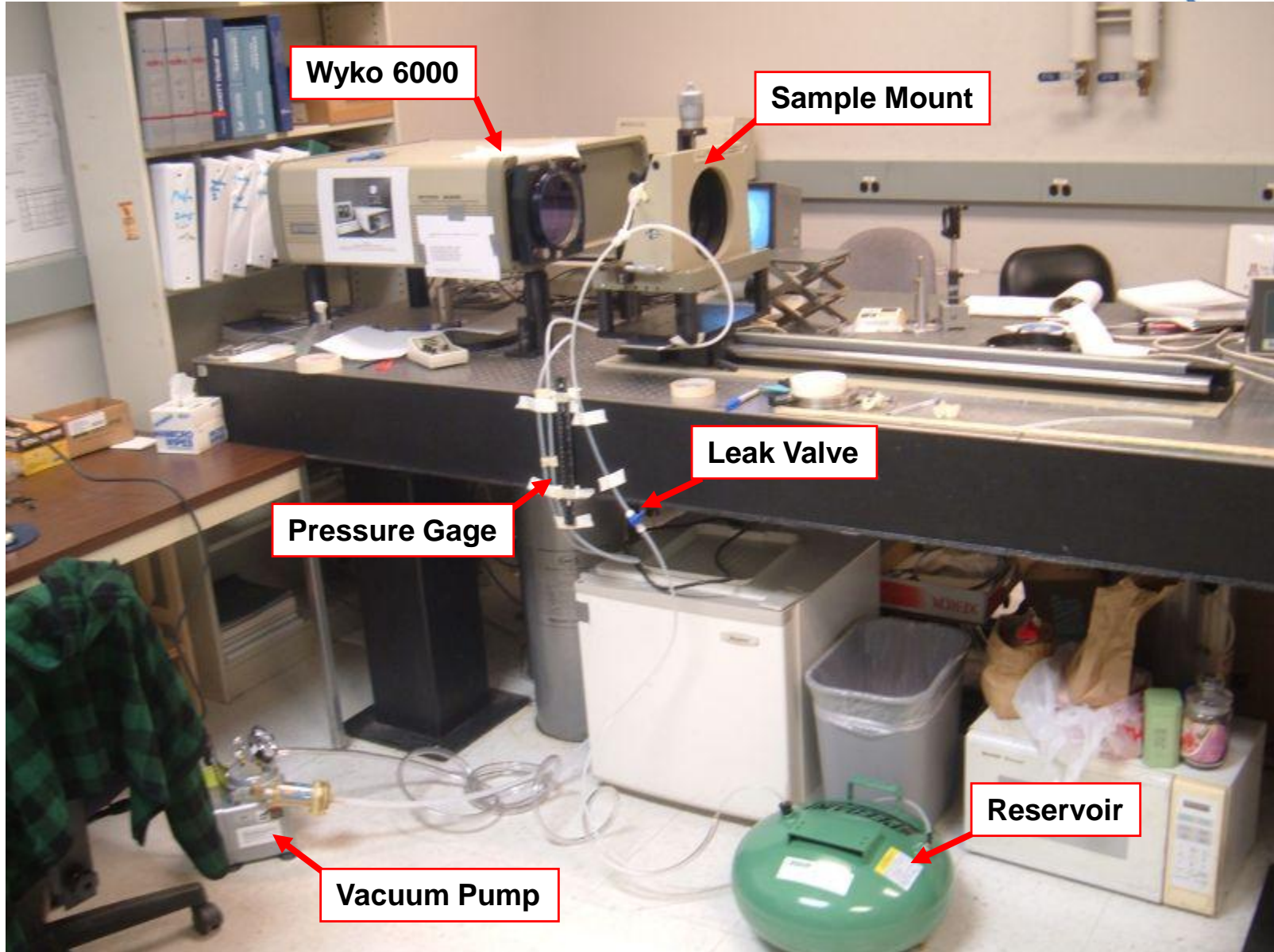
	Requirements		Verifications
1.	Difference between simulation and experiment	< 20 [%]	Test
2.	Supported with edge face	NA	Inspection
3.	Surface Slope Irregularity	< 10 [nm/cm RMS]	Analysis
4.	Resonant Frequency	> 30 [Hz]	Analysis
5.	Stability - power - w/o power	20 [nm] 3 [nm]	Analysis

# Experimental Setup 1

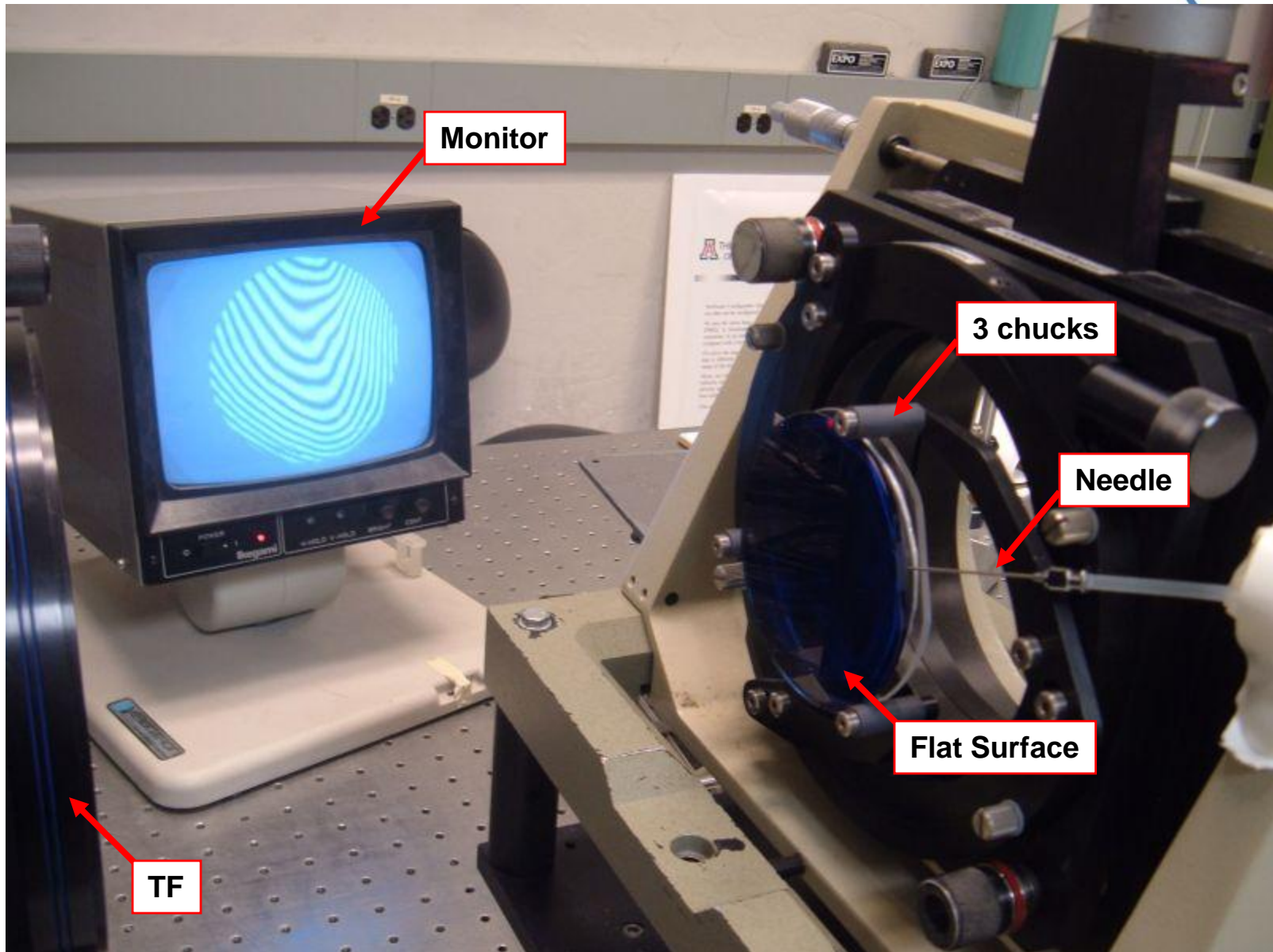


Sensitive around air pressure  
Cheap

# Experimental Setup 2

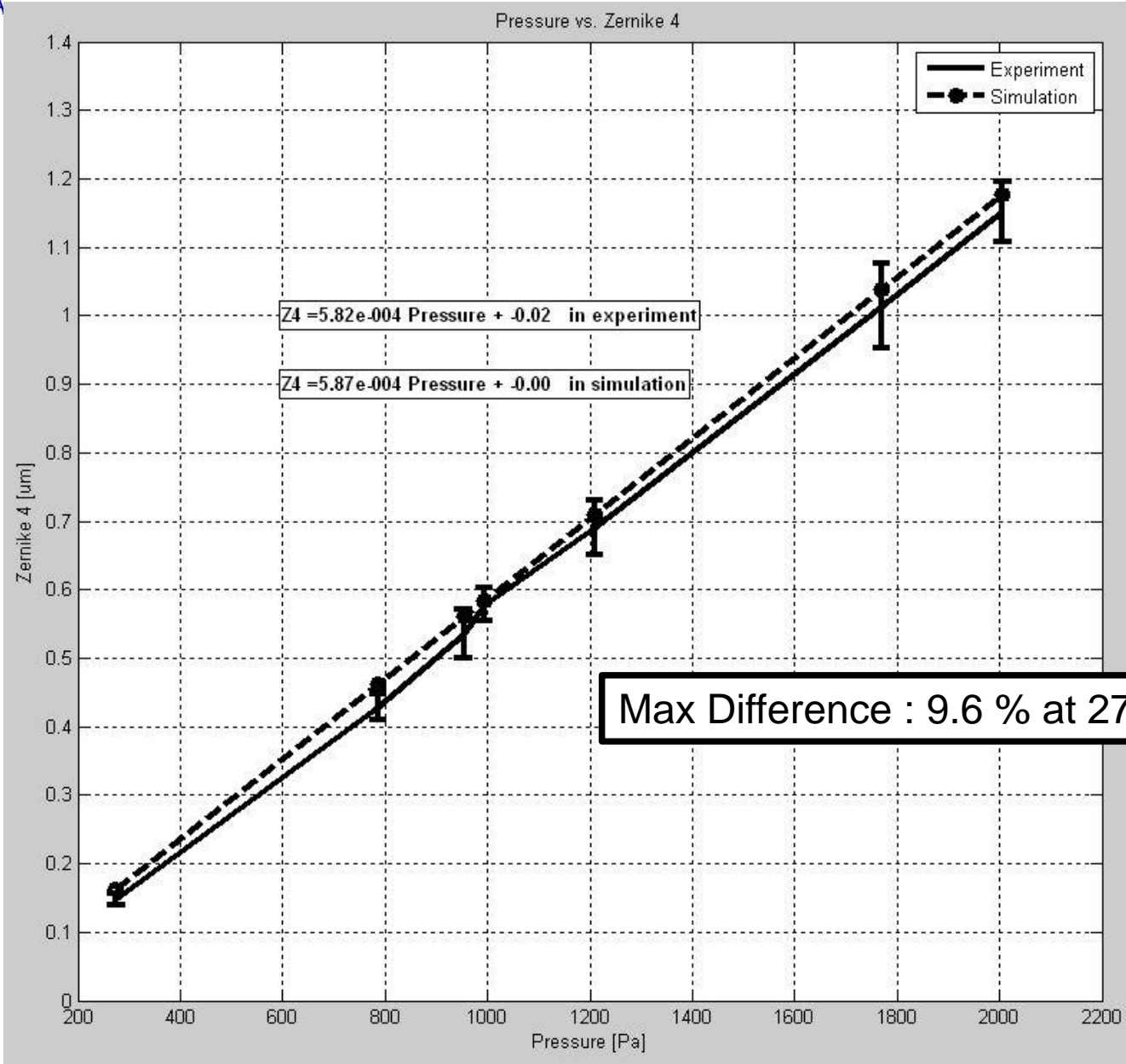


# Experimental Setup 3

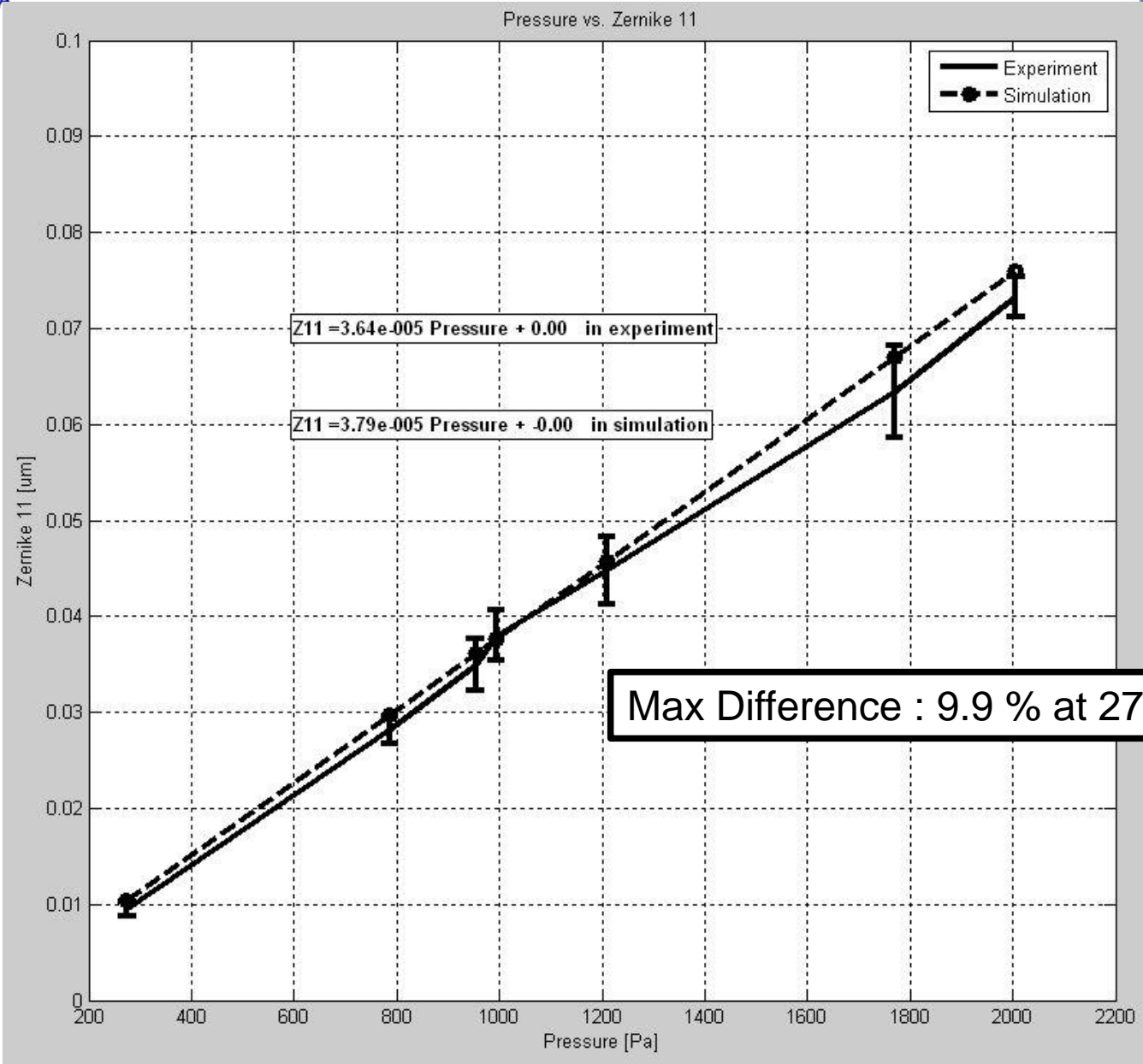




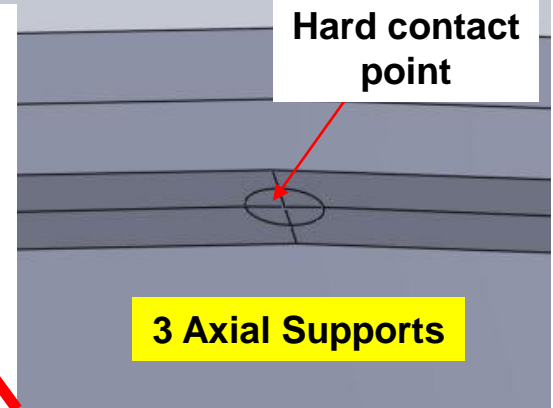
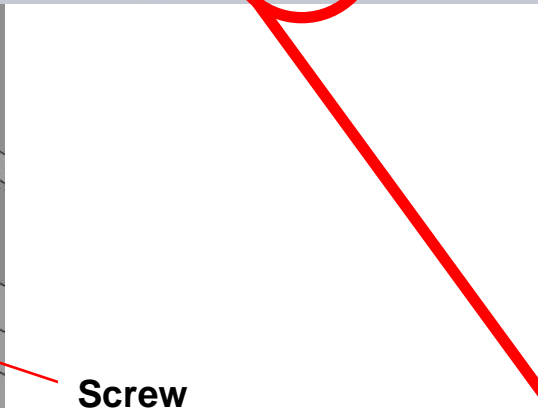
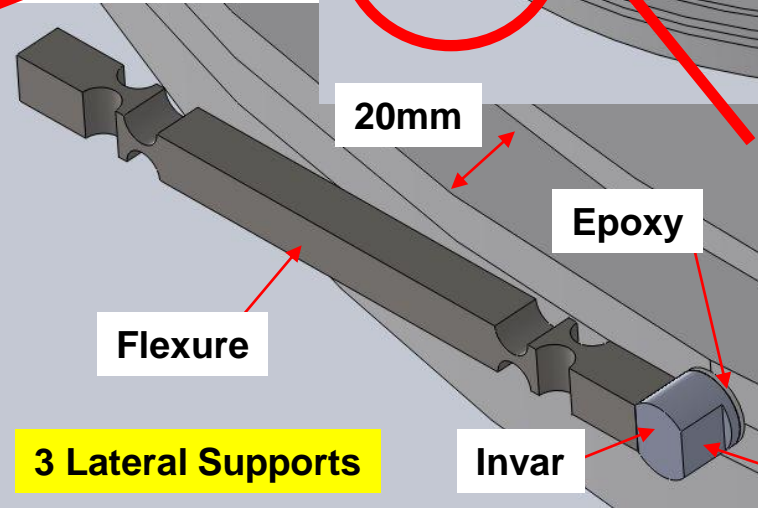
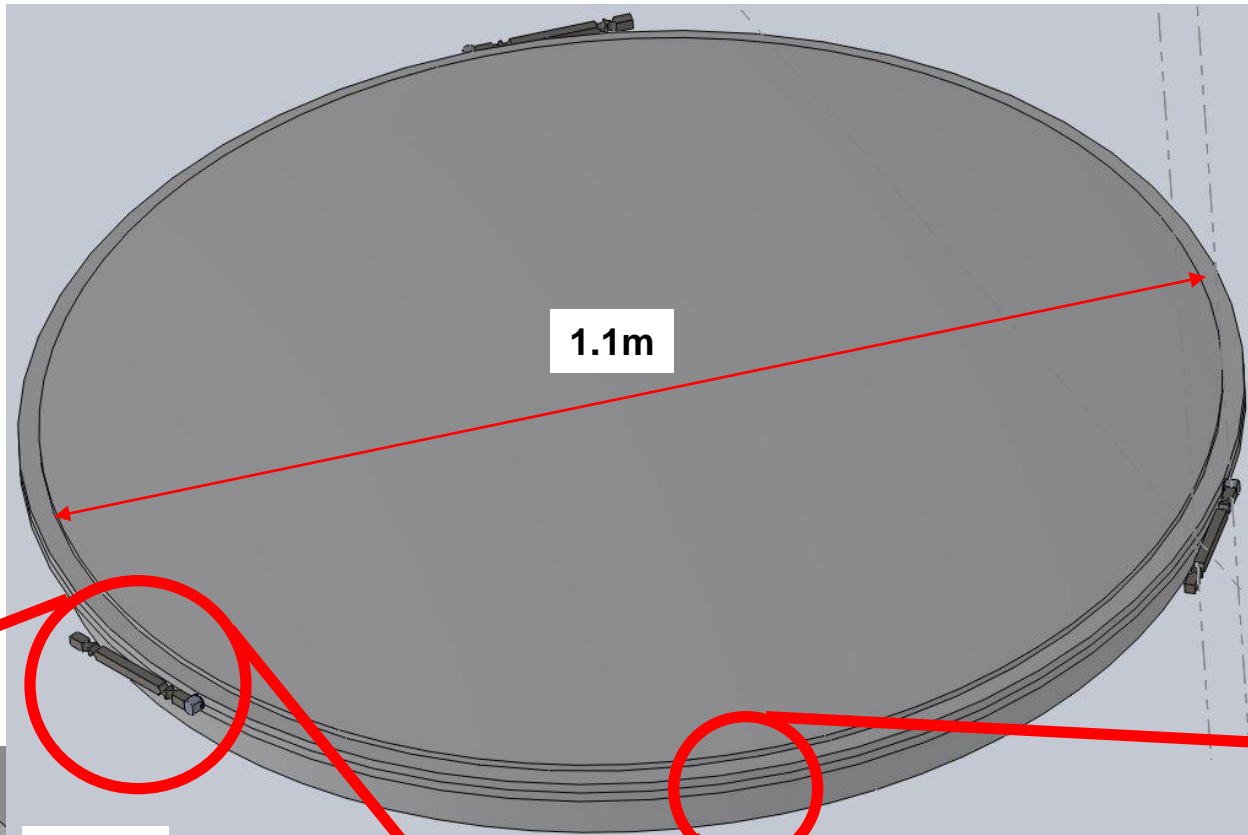
# Experimental Results (Power; Z4)



# Experimental Results ( $\rho^4$ term; Z11)

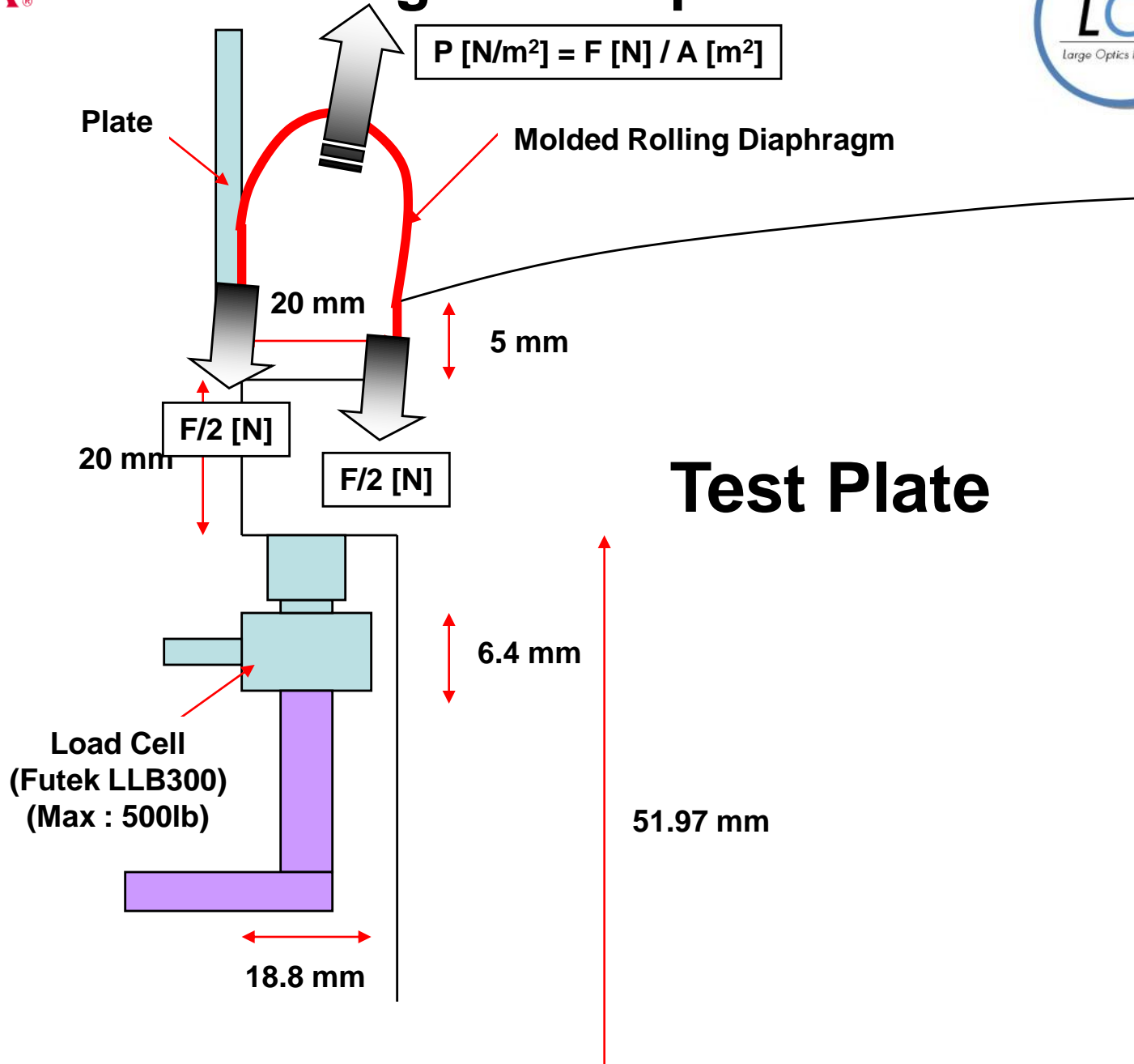


# Design Concept 1

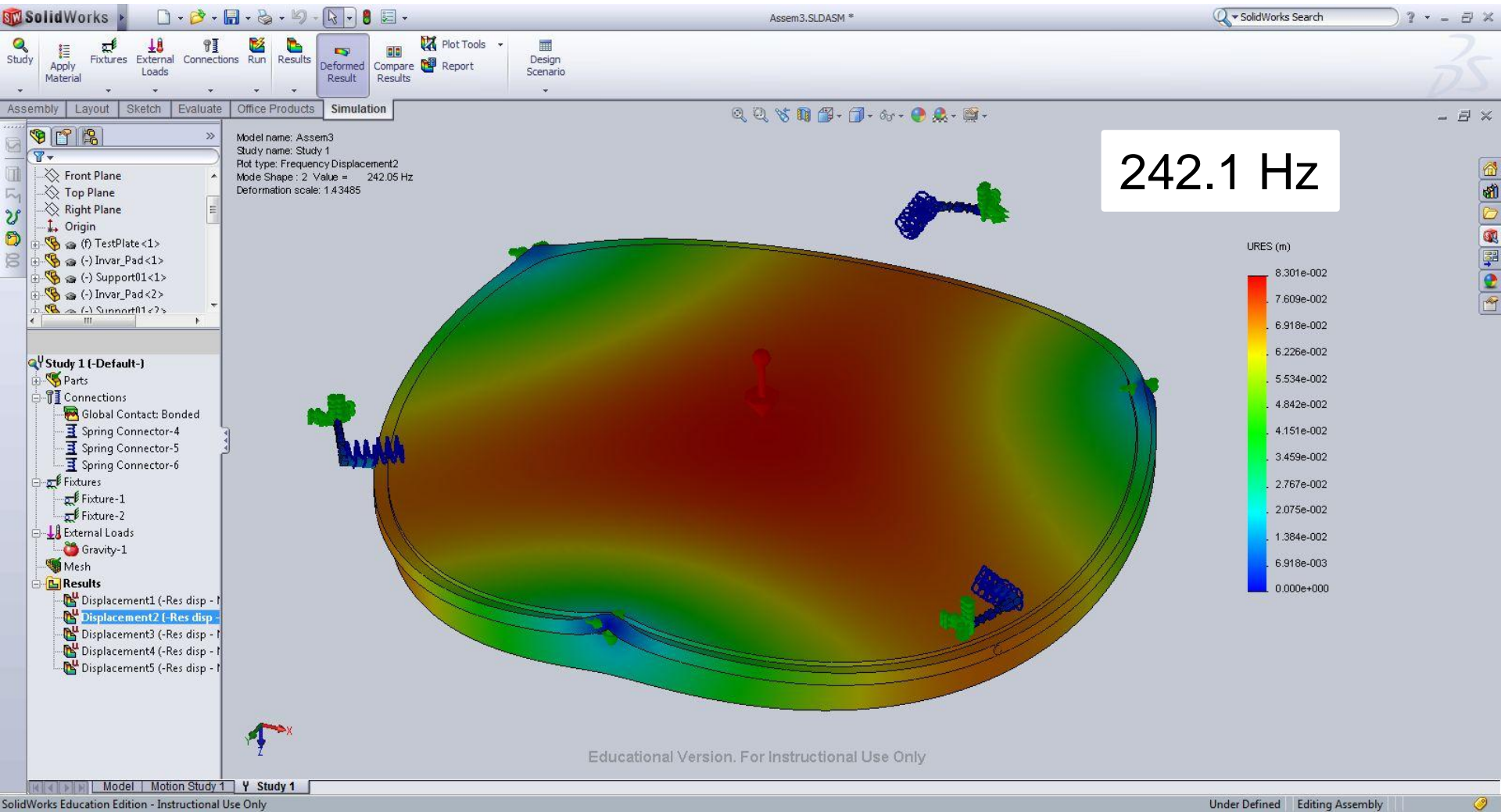


# Design Concept 2

$$P \text{ [N/m}^2\text{]} = F \text{ [N]} / A \text{ [m}^2\text{]}$$



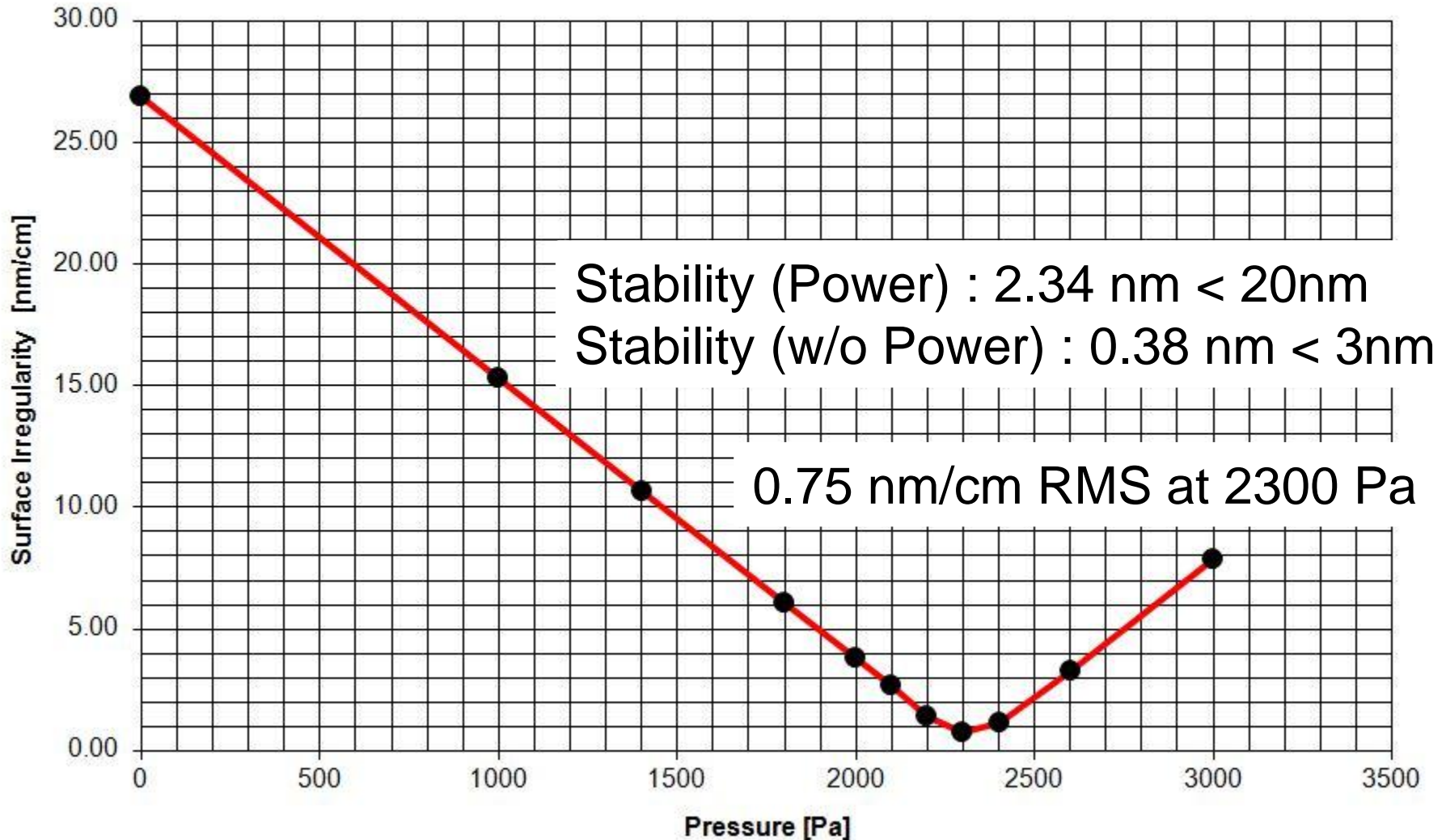
# Resonant Frequency



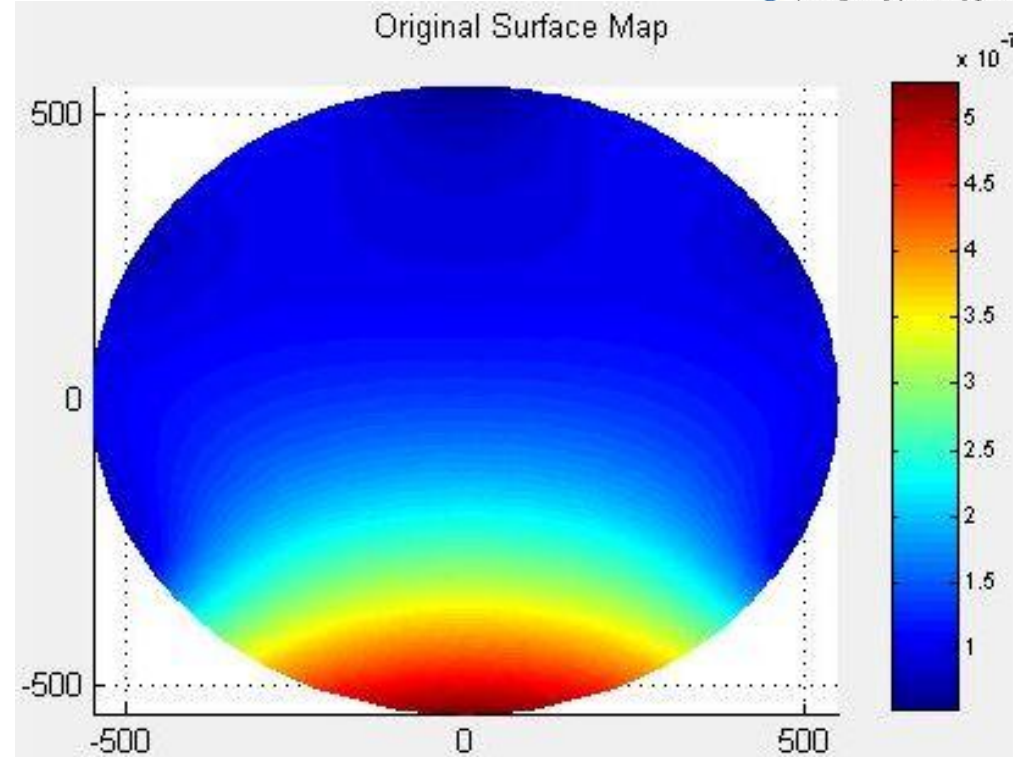
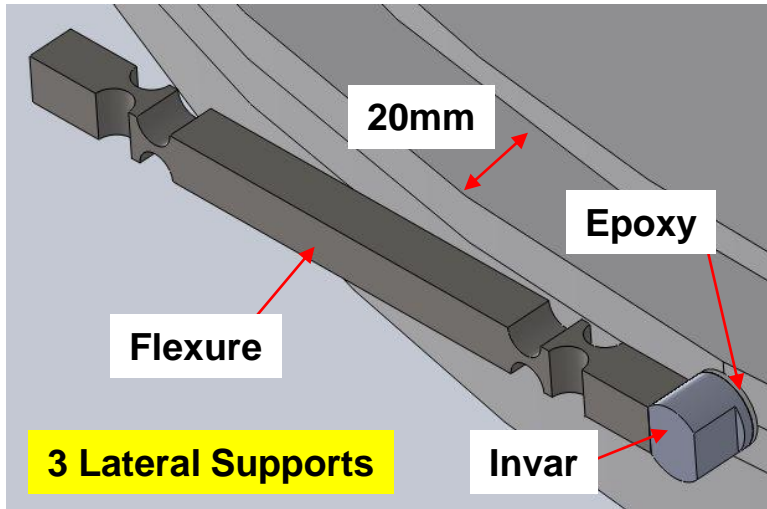
Educational Version. For Instructional Use Only

# Pressure vs. Surface Slope Irregularity

Pressure vs. Surface Slope Irregularity of Test Plate



# Flexure Effect



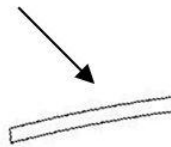
(Matlab code by Won Hyun Park)

100 um tolerance at flexure causes 0.2 N-m moment at the test plate.  
 This causes 2.71 nm/cm RMS surface slope irregularity.

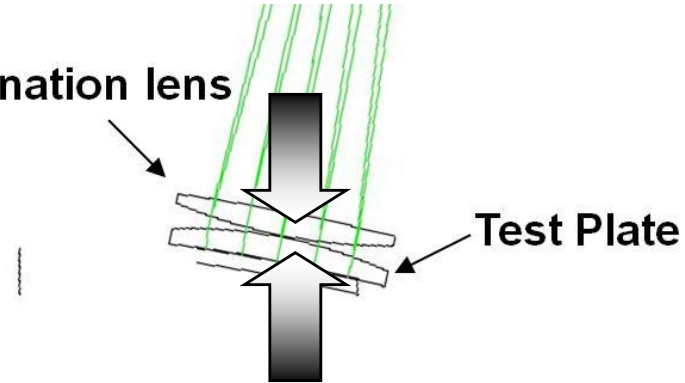
Since this issue is independent from vacuum support,  
 (Total RMS) =  $\text{SQRT}((0.75 \text{ nm/cm})^2 + 3 \times (2.71 \text{ nm/cm})^2) = 4.75 \text{ nm/cm RMS}$   
 $< 10 \text{ nm/cm RMS}$

# Interferogram with Vacuum Support

Secondary Mirror



Illumination lens

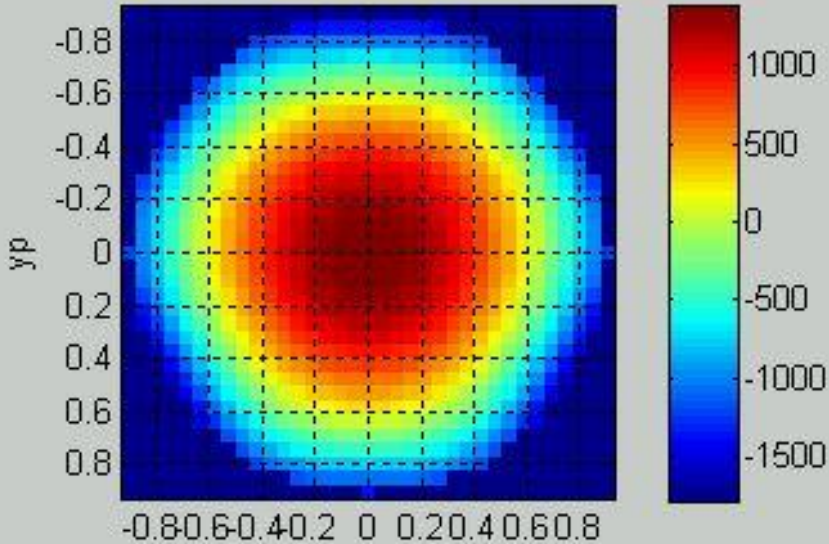


Test Plate

Piston and Tilts are removed.

RMS = 809.3 (nm)

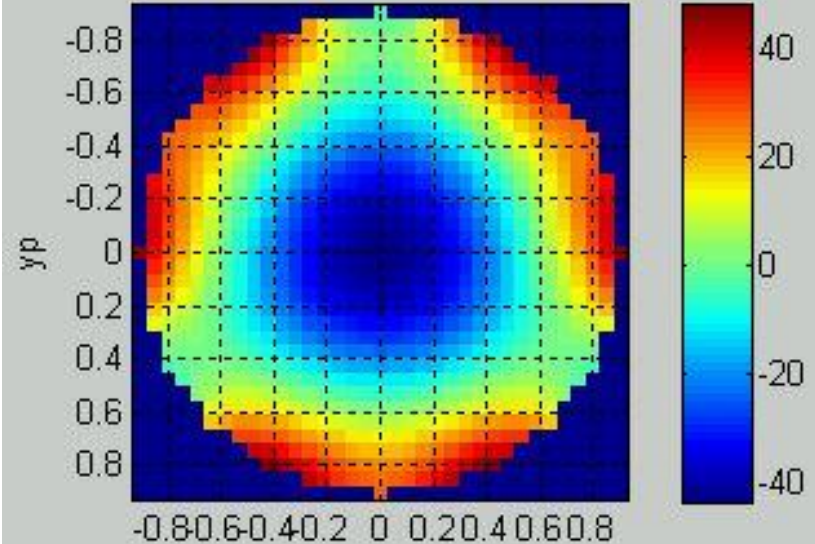
OPD in nm



W/O Vacuum Support

RMS = 22.05 (nm)

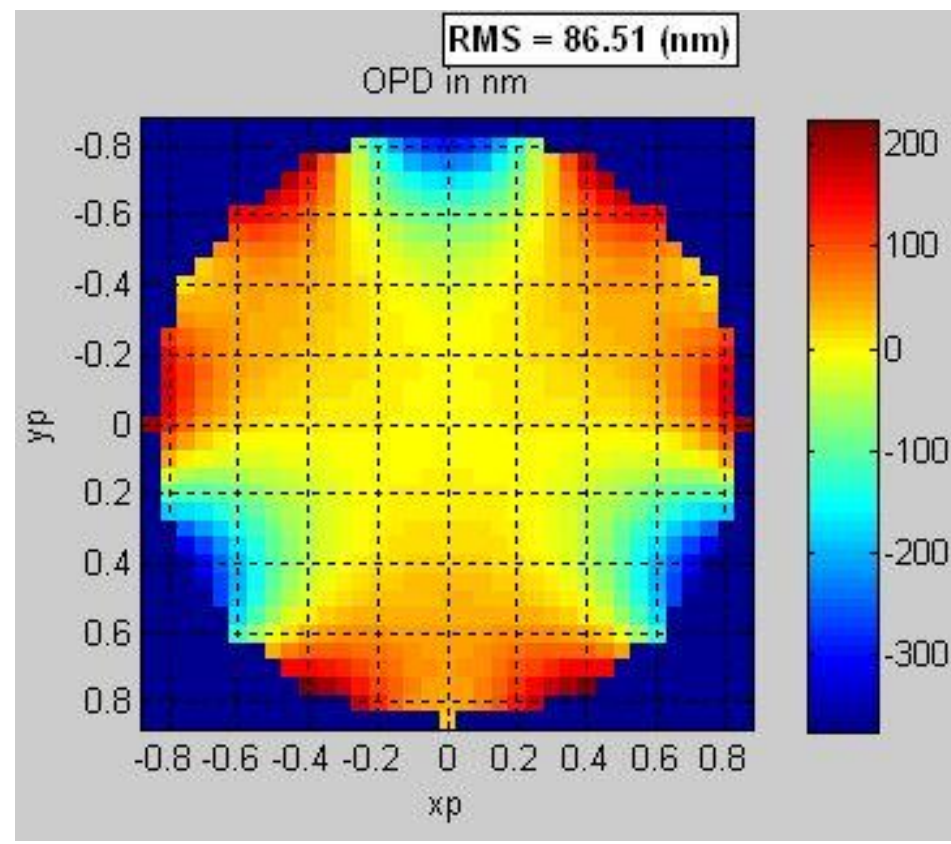
OPD in nm



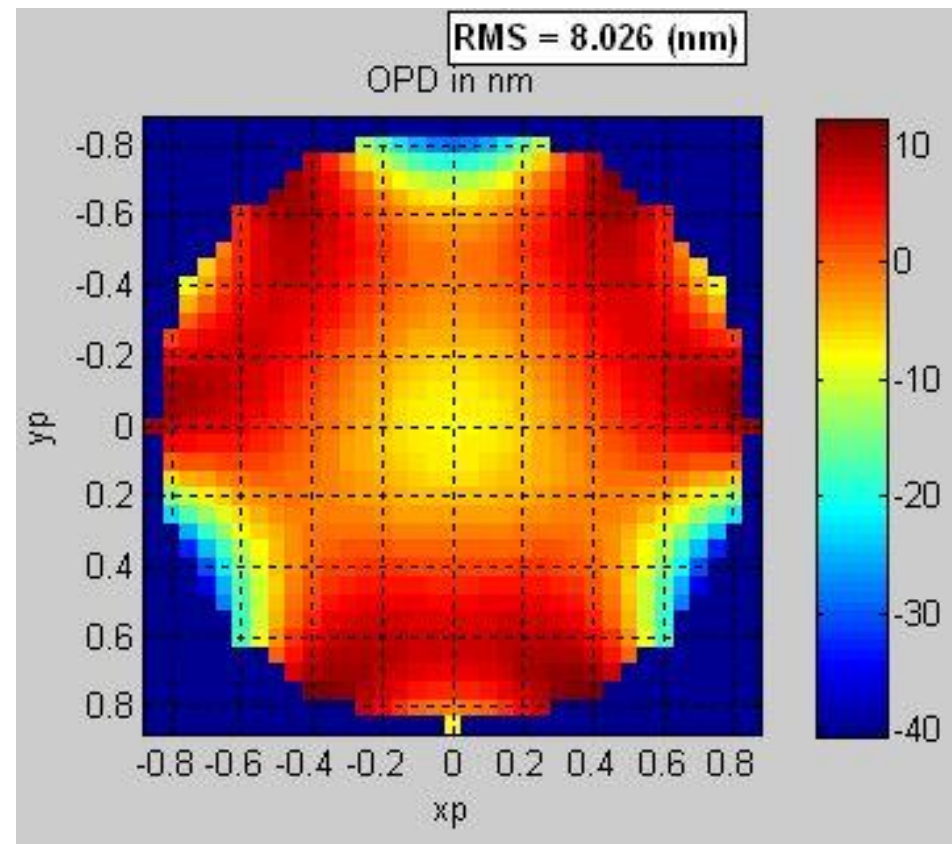
W/ Vacuum Support



Piston, Tilts, and Power are removed.



W/O Vacuum Support



W/ Vacuum Support

# Summary



- I did the simple experiment and the simulation to evaluate the vacuum support.
  - Difference between exp. and sim. was  $10\% < 20\%$  in spec.
  - Surface slope irregularity was  $4.75 \text{ nm/cm} < 10 \text{ nm/cm}$  in spec.
  - Resonant Frequency was  $242.1 \text{ Hz} > 30 \text{ Hz}$  in spec.
  - Stability of Power was  $2.34 \text{ nm} < 20 \text{ nm}$
  - Stability of w/o Power was  $0.38 \text{ nm} < 3 \text{ nm}$
- The vacuum support for the large interferometric reference surface can be effective.
- Future Work
  - Detail design around the diaphragm
  - Check an effect of index change
  - Comparison with actual experiment